NGSLR:
NASA’s Next Generation
Satellite Laser Ranging System

Jan McGarry
NASA/GSFC/694
Solar System Exploration Division

September 27, 2011
Overview

NGSLR is a high repetition rate single photon detection laser ranging system capable of tracking cube corner reflector (CCR) equipped satellites in Earth orbit. The concept of NGSLR was developed by J. Degnan (GSFC, retired) in the 1990s. Technical development continues at Goddard. The system has demonstrated tracking of Earth orbit satellites with altitudes from ~ 1000 km to 20000 km. Completion of the NGSLR prototype will occur during the Space Geodesy Proposal.

System Features:
- 1 to 2 arcsecond pointing/tracking accuracy
- Track CCR equipped satellites to 20,000 km altitude, 24/7 operation
- Reduced ocular, chemical, electrical hazards
- Semi automated tracking features
- Small, compact, low maintenance, increased reliability
- Lower operating/replication costs
Team

**John Annen** (automation lab lead, hardware development) - GSFC/693
**Jack Cheek** (software lead, sysadmin, real-time software) - SigmaSpace
**Bart Clarke** (software: OD, SLR data analysis, signal processing) - Honeywell
**John Degnan** (lasers, SLR theory, original SLR2000 concept) - SigmaSpace
**Bud Donovan** (hardware lead, deputy NGSLR lead, laser safety) – Honeywell
**Felipe Hall** (documentation lead) – Honeywell
**Evan Hoffman** (hardware development & sustaining engineering) - Honeywell
**Julie Horvath** (config control & collocation lead, i/f with ILRS) - Honeywell
**Tony Mann** (software: unix sysadmin, device drivers) - Honeywell
**Jan McGarry** (NGSLR lead, algorithm & software development) - GSFC/694
**Don Patterson** (hardware develop., sustaining & system engineering) - Honeywell
**Randy Ricklefs** (software & OS, device drivers, algorithm) – Cybioms
**Mark Torrence** (SLR analysis) - SGT
**Tom Varghese** (system engineering) – Cybioms
**Scott Wetzel** (cross-SGP integration) - Honeywell
**Tom Zagwodzki** (system engineering) - Cybioms

*Combined total of > 250 years SLR experience*
Historically Important Dates
NGSLR was originally called SLR2000

1994: Original SLR2000 concept, limited to LAGEOS (< 8500 km slant range) and lower satellites, was first presented to, and endorsed by, the NASA Belmont SLR Workshop.
1997: SLR2000 technical approach reviewed and approved by GSFC MTPE Office (Dr. Robert Price). First substantial funding provided.
2000: Revised 2000 ANSI laser safety standards for subnanosecond pulses have major impact on maximum transmitted energy and minimum pulsewidth affecting both receive signal strength and single shot range accuracy.
2001: All major subsystem tests completed successfully.
2002: Prototype system assembled and ready for field tests. Phase III laser on order from Q-Peak.
2003: Visual tracking of sunlit satellites demonstrated. Two to three arcsecond star calibrations routinely achieved.
2004: LEO satellite tracking demonstrated under operator control.
2006: Many automated subsystems (variable beam expander, PRF, upgraded sky and star cameras) are integrated and operational except for transmitter point ahead and quadrant detector pointing correction.
2009: Robust day/night tracking of LAGEOS and night tracks of GNSS satellites (GLONASS, Etalon) with eyesafe laser and transmitter point-ahead capability demonstrates original goals of SLR2000 development.
2011: SGP starts its 2-year effort, major funding received.
Original Goals*

“...the primary technical goals of the SLR 2000 system are:

Unmanned, eyesafe operation
24 hour tracking of LAGEOS and lower satellites
One centimeter (RMS) single shot precision or better
Minimum 100 ranges per normal point
Mean time between failures: > 4 months
Automated two-way communications with a central data processor via Internet
System free of optical, electrical, and chemical hazards

Secondary goals for the system, presently viewed as highly desirable but perhaps difficult to achieve, include a capability to range to high altitude satellites such as GPS, GLONASS, and ETALON and the ability to retrofit two color technology at some later date.”

Current Requirements

- 24 hour tracking of LEO, LAGEOS & GNSS satellites (that have ILRS approved retro-reflector lidar cross sections)
- One millimeter normal point precision on LAGEOS
- Accuracy and stability at the MOBLAS level or better
- Semi-autonomous operations
- Radar for all satellite ranging
- Mean time between failures: > 4 months
SYSTEM Description
## Major Subsystems

1. Time & Frequency
2. Telescope
3. Transceiver Bench
4. Laser
5. Laser Hazard Reduction System (LHRS)
6. Tracking
7. Receiver
8. Computer and Software
9. Weather
10. Shelter and Dome
NGSLR System Characteristics

• **Telescope:**
  - 40 cm Telescope Aperture Off-Axis Parabola
  - No Central Obscuration

• **Tracking:**
  - AZ/EL with 1 arcsec RMS gimbal pointing accuracy

• **Transceiver Bench:**
  - Common Optics for Transmit and Receive
  - Passive Transmit/Receive Switch
  - Risley Prism Point-Ahead of Transmit

• **Laser:**
  - Subnanosecond pulse, 2 kHz
  - Asynchronous PRF, software controlled
  - Divergence control by software

• **Receiver:**
  - High QE, GaAsP Microchannel Plate Photomultiplier
  - Constant Fraction Discriminators
  - GPS-synchronized Rubidium Oscillator /Time and Frequency Receiver
  - Picosecond Precision Event Timer

• **Weather:**
  - Day/Night All-Sky Cloud Sensor (thermal)
  - Wind Monitor
  - Surface Pressure, Temperature, and Humidity Monitors
  - Visibility/Precipitation Sensor
NGSLR System Block Diagram

Simplified Hardware Block Diagram

This diagram shows the basic flow of signals within the system.
Software/Computer Overview

▪ Computers:
  - Pseudo Operator (POP) – performs operator decisions
  - Device Access Manager (DAM) – optical bench controller
  - Interface & Control Computer (ICC) – real-time data I/O
  - Remote Access Terminal (RAT) – interface to human
  - Analysis Computer (ANA) – post-processing
  - Camera computer – start and sky camera interface
  - Dome controller – slaves dome to telescope during operations

▪ Backplanes: VME, PCI, ISA.

▪ Software
  - Operating systems: LynxOS, Linux, Windows, DOS
  - Languages: “C”, assembly, perl.
  - Lines of code: ~200,000
NGSLR Computer Interfaces

* Shared memory is accomplished using the Bit-3 PCI to VMEbus adapter, which includes 8MB of RAM.

Key
- Data Connection
- Ethernet
- Remote Connectivity

Remote Access Terminal (RAT)
- RAT is connected to Ratsnest on DAM via Ethernet

Analysis Computer (ANA)
- (Post Processing Analysis)

Central Facility

Camera computer

Dome Controller

I/O Chassis
- The I/O Chassis controls beam blocks, the radar, and other components.

Timing

ICC

POP

DAM

Ranging Electronics

Mount

Laser

Weather

Transceiver Bench

StarCamera

SkyCamera

USB

FireWire
Computer & Equipment Racks

Cesium Frequency Standard Distribution Amplifier
NASA Radar Controller
GPS Time & Frequency Receiver
Computer Clock Sync Interface (Timing Box)
Dome Controller
Dome Motor Driver
Dome Shutter Controller
Laser Block "Enable"
Interface Control Computer (ICC)
Xybion Mount Controller
UPS

HP Laser Fire Frequency Counter

Power supply for the video camera on the telescope
I/O Chassis
Event Timer (ET)
Range Gate Generator (RGG)
Camera Computer

VME Backplane with:
- POP
- DAM
- Shared Memory Card
- Interface Cards
(See R-02 for details on the VME Chassis)

Left: Axis Video Server
Right: Ethernet Switch

Rack 1
Rack 2

UPS
Automation Overview
(some of the functions that will be performed by software)

• Obtaining input files:
  - automatically pull prediction and other data files from the server.

• System scheduling:
  - software completely determines/controls what is tracked and when.

• Operator decision making:
  - open/close dome based on weather,
  - keep telescope from pointing into the sun,
  - determine if we can track and where in the sky based on cloud cover.

• Signal processing and closed-loop tracking:
  - determine if system is hitting the satellite,
  - search for the satellite and optimize the pointing.

• Transmit / receive path optics configuration and control:
  - determine and control optical bench configuration,
  - decide configuration based on target, day/night.

• Data processing and product delivery: normal points delivered hourly.
New Technologies Developed for NGSLR

The requirements of SLR2000 (i.e. eye safety and unmanned operation) led to a number of unique computer-controlled hardware devices including:

• **Totally Passive Transmit/Receive Switch** allows the full aperture of the telescope to be shared simultaneously, with minimal optical loss, by the transmitter (for eye safety) and receiver (for signal strength) independent of the laser repetition rate and receive signal polarization.

• **Transmitter Beam Expander** allows the transmitter beam divergence to be varied as a function of satellite range for enhanced signal strength while maintaining a fixed beam diameter at the telescope exit window for eye safety.

• **Variable Spectral Filter** optimizes the filter transmission and spectral bandwidth for daylight, twilight, and night operations.

• **Dual Risley Prism Device** permits independent arcsecond accuracy pointing of the transmitter and receiver allowing smaller receiver fields-of-view for reduced solar noise.

• **Variable Iris Spatial Filter** allows adjustment of the receive FOV for less solar noise.
New Technologies Developed for NGSLR (continued)

• **Variable Laser Trigger** varies laser repetition rate about the nominal 2 kHz to prevent backscatter from the outgoing laser pulse from overlapping satellite returns at receiver.

• **Dual Liquid Crystal Optical Gates** further reduces laser instrument and atmospheric backscatter by more than two orders of magnitude, independent of polarization.

• **Smart Meteorological Station** monitors hemispherical cloud cover and ground visibility (to support satellite selection and efficient operations), precipitation (for system protection), wind speed and direction, while providing the usual atmospheric surface pressure, temperature, and relative humidity measurements needed to support atmospheric refraction corrections to the range measurements.

• **Algorithms and related software** to give NGSLR the ability to autonomously (1) determine when and how to change the laser pulse repetition frequency (PRF) to avoid collision between outgoing and incoming laser pulses, (2) process returns to find satellite events in very low signal to noise environments, (3) continually monitor the angular proximity of the telescope to the sun and move the mount to avoid getting sunlight into the detector, (4) direct the transmit beam as an angular offset from the telescope to put laser pulses where the satellite will be when the light arrives.
Original System Configuration
Q-Peak Laser and Photek QMCP

Q-Peak Laser:
• Model MPV-2000, doubled YVO$_4$
• 200 µJ, 532 nm, 350 ps, 2 kHz, 0.9 x 1.8 mR, 1.1 x 0.5 mm Ø
• Rationale for Replacement
  - Laser performance degraded – energy decreased from 200 µJ to 60 µJ
  - Daytime GNSS tracking requirement necessitated increased energy

Photek QMCP Installation:
• Quadrant Anode, 12% QE, 140 ps $t_r$
• Rationale for Replacement
  - QE decreased from 12% to <6%
  - New technology MCPs have High QE >30%

SGP 9/27/2011 jlfm
NGSLR Performance with Q-Peak Laser

- **Q-Peak Laser**
  - Low energy/eye-safe
  - 2KHz repetition rate
  - 350 ps pulsewidth

- **NGSLR Successful Tracking with Q-Peak Laser**
  - LEO (day/night):
    - BEC, Jason, Ajisai, Starlette, Stella
  - Lageos 1 & 2 (day/night)
  - HEO (night)
    - Etalon, GLONASS

- **NGSLR System Performance**
  - Ground Cal RMS: 20mm
  - Lageos Single Shot RMS: 40 mm
  - Lageos Normal Point RMS: mm’s
  - Ground cal stability: +/- 5mm
  - Bias between NGSLR and MOB-7: ~2 cm
Current Automation Status

- **Completed, tested and working**
  - Automated star calibration
  - Sun avoidance (software)
  - Laser pulse collision avoidance
  - Point-ahead tracking using Risley Prisms
  - Obtaining input data and system scheduling

- **Almost complete**
  - Open/close of dome based on weather
  - Determination what can be tracked based on cloud cover
  - Determination if system is hitting the satellite
  - Search for the satellite if not getting signal returns
  - Normal point generation and data delivery

- **Being worked**
  - Remaining optical bench motor controls
  - New I/O chassis software interface
  - Closed loop tracking (optimizing biases during tracking)
  - Fully automated ground calibration
NGSLR Documentation

- Production of Documentation is proceeding as planned
- All critical documentation will be updated before December 2012, documenting the NGSLR system in its final state.

NGSLR Documentation Schedule

<table>
<thead>
<tr>
<th>Document Number</th>
<th>Title</th>
<th>Year</th>
<th>Description</th>
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<tbody>
<tr>
<td>NASA-NGSLR-DWG-06</td>
<td>NGSLR - System Drawing Set</td>
<td>4/1/2011</td>
<td>Preliminary version completed</td>
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<tr>
<td>NASA-NGSLR-OPS-NGSLR(v0.6)</td>
<td>NGSLR - Operations Manual (EyeSafe)</td>
<td>7/1/2011</td>
<td>Preliminary version completed</td>
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<tr>
<td>NASA-NGSLR-DWG-07</td>
<td>NGSLR - Engineering Drawings for Automation of Optical Bench</td>
<td>10/1/2011</td>
<td>Projected completion date</td>
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<tr>
<td>NASA-NGSLR-HWR-Dome</td>
<td>NGSLR - Dome Guide</td>
<td>1/1/2012</td>
<td>Projected completion date</td>
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<tr>
<td>NASA-NGSLR-SWR-Camera</td>
<td>NGSLR - Star Calibration and Sky Camera Software Guide</td>
<td>5/1/2012</td>
<td>Projected completion date</td>
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<tr>
<td>NASA-NGSLR-HWR-SysInterface</td>
<td>NGSLR - System Interface Manual and Drawing Set</td>
<td>7/1/2012</td>
<td>Projected completion date</td>
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<tr>
<td>NASA-NGSLR-INT-Overview</td>
<td>NGSLR - General Systems Overview</td>
<td>8/1/2012</td>
<td>Projected completion date</td>
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Bring all manuals up to Version 1.0 12/1/2012 Update all manuals

Post 2012

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<tr>
<td>NASA-NGSLR-HWR-Install</td>
<td>NGSLR - Hardware Installation Manual</td>
<td>TBD</td>
<td>Date indicates a preliminary revision</td>
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<tr>
<td>NASA-NGSLR-SWR-Scripts</td>
<td>NGSLR - Scheduling and other System Scripts</td>
<td>TBD</td>
<td>Date indicates a preliminary revision</td>
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<tr>
<td>NASA-NGSLR-SWR-Sim.Manual</td>
<td>NGSLR - Simulation Software Manual (ICC, RAT, POP, DAN)</td>
<td>TBD</td>
<td>Date indicates a preliminary revision</td>
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<tr>
<td>NASA-NGSLR-SWR-Sys_Admin</td>
<td>NGSLR - System Administration &amp; Maintenance Manual</td>
<td>TBD</td>
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The Path to Completion
Configuration for Intercomparison with MOB-7

- **Laser and detector:**
  - NASA in-house built laser (Poulios & Coyle): ~1 mJ per pulse, 200 ps pulsewidth
  - Variable rep rate (nominally 2 kHz), variable transmit energy
  - Hamamatsu high QE MCP detector
  - Expect a factor of 20x improvement over Q-Peak configuration

- **Tracking performance for period Dec 2010 – July 2011:**
  - Successfully tracked LEOs, LAGEOS and GLONASS at night
  - Return rates were higher than with Q-Peak
  - Was not able to range during daylight – due to wavelength instability and mismatch with daylight filter

- **In-house laser issues currently being worked in B33 lab:**
  - Internal damage from high energy density on laser optics (this has been fixed)
  - Stability issues (purchased new seeder to eliminate wavelength instability)
  - Comes back shortly to NGSLR to allow us to continue development efforts and performance improvements, and to establish a baseline for system tracking / ranging performance
In-house built 2 kHz laser
Regenerative amplifier seeded by a gain-switched diode laser

Nd:YAG slabs
1.5m folded cavity
Seeder

KTP doubler
Pockels cell
External Cal Data Session Stability  Plot of External Cal Residual (millimeter) and Data Rate (%) vs. Time for a duration of 75 minutes; 4\(^{th}\) channel (BLACK line) in the upper plot shows the mean of 3 channels; Lower plot shows the data rates for each of the Q1 thru Q4 Channels; Horizontal scale is 5 minutes
NGSLR Automation Laboratory

• Development Efforts in Lab
  - Development of new optical bench layout
  - Design and testing of specialized mounting hardware for optics
  - Verification of selected optical components
  - Design and verification of alignment aids and alignment procedure
  - Assembly and alignment of new optical bench
  - Verification and Testing of COTS high energy, short pulse laser
  - Integration of COTS laser onto optical bench

• Benefits of Lab Work
  - Introduce a fully automated layout
  - Isolate transmit/receive path on the bench to reduce system noise
  - Improve access to all devices for alignment
  - Upgrade alignment procedure
  - Increase system efficiency
  - Allows work to go on without disturbing NGSLR tracking

Once aligned and tested, the new bench will be moved into the NGSLR shelter, accelerating the integration process
New Optical Bench
preliminary design
IO Chassis

- Controls and distributes the proper gate signals for the PMT and discriminator during ground calibrations and satellite tracking using its internal delay circuits and the RGG inputs as well as electronics for the MCP blanking circuit.

- Contains electronics and firmware enabling computer interface and control and/or monitoring of various hardware on the Optical Bench such as beam blocks, optical density filters, shutters, gradient ND filter wheels, etc.

- Provides interface electronics for the Remote Control Box which allows manual control of various elements on the Optical Bench for maintenance and alignment purposes.

- Serves as the safety interlock chassis providing power and control to beam blocks when a safety condition occurs such as an aircraft detect, opening of the Shelter door, and unauthorized access of the stairway to the telescope area.

- Provides the interface electronics for the control of the radar subsystem.
Software Development and Testing (Using NGSLR Automation Lab)

- Facilitates the development of new software without disturbing the operational system and serves as a test location for operational spare computers.
- Uses a complete set of duplicate computers, motor controllers, an optical bench, and an IO chassis (both a software simulator and a hardware setup).
- Uses software simulators when/where spare components are not available.
- Enables the designing, coding and testing of IO chassis software and automation control software.
- Permits testing of the automated search and cloud decision software.
- Allows the testing of star assessments, star calibrations, ground calibrations, satellite tracking as well as “real world” scenarios.
Final System Configuration

Hamamatsu High QE MCP PMT:
- Model R5916U-64
- GaAsP Photocathode
- QE > 43%
- Rise Time < 178 ps
- Photocathode Input = 25mm Ø

Photonics Industries Short Pulse, Hi Energy, Hi Rep Rate Laser:
- Model RGL532-2.5
- Maximum Energy = 3 mJ
- Pulse Width FWHM = 50 ps
- Repetition Rate = Single Shot to 5 kHz
- Beam Divergence < 1 mR
- Output Beam Diameter = 1.7mm
- Spatial Mode Profile = TEM\(_{00}\)
- Long Term Stability < +/- 2%
- Pulse to Pulse Stability < 2% RMS

Laser will not be available for use until early 2012
NGSLR Collocation with MOBLAS-7

• What is a Collocation?
  - Collocation is the process of geometrically comparing ranging data from two or more SLR systems in close proximity (preferably <60 meters) by simultaneously ranging to common targets, ground and satellite based.

• Why do Collocation?
  - Collocation analysis is the best engineering tool available for rapid identification of sub-centimeter systematic error sources.
  - Because the Collocation technique was used extensively prior to deploying the Heritage NASA Network, it is probably the reason that the NASA SLR systems have performed with so few biases since the mid-1980’s. The NASA SLR systems are still considered some of the best in the world.

• NGSLR / MOBLAS-7 Collocation will establish criteria for, and verify:
  - System Stability
  - System Precision
  - System Accuracy
  - System Timing
  - System Bias
SUMMARY

- Most of the original goals for SLR2000 have been achieved.
- Significant progress has been made in performance and automation.
- LEO to LAGEOS ranging demonstrated day & night with eye-safe energies.
- GNSS tracking demonstrated at night with eye-safe energies.
- Major technology developments:
  - Passive T/R switch, Risleys for point-ahead, transmitter beam expander,
  - smart Met station, liquid crystal optical gates, ...
- Major software achievements:
  - Sun avoidance software, laser PRF changes for pulse collision avoidance,
  - point-ahead angular calculations, signal processing, operator decision
  - automation, ...
- NGSLR lab has been setup - allowing parallel progress in automation,
  - software checkout and system performance.
- Significant progress on documentation – have initial version of several
  - major system documents & drawings.

Most importantly, NGSLR has:
- Achievable path to completion, and
- Accomplished team with extensive SLR experience.