

# Toward high-rate on-time mm-accurate SLR at Stafford, Virginia

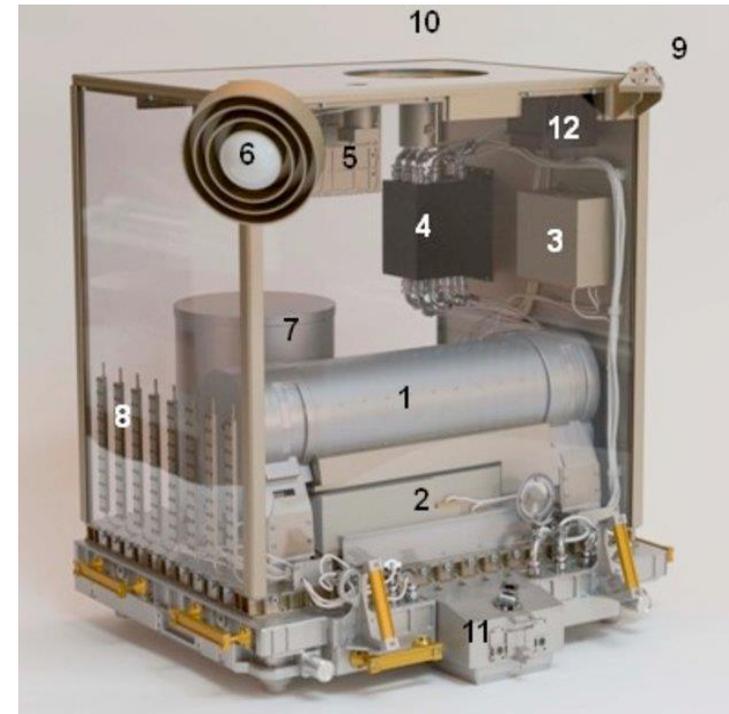
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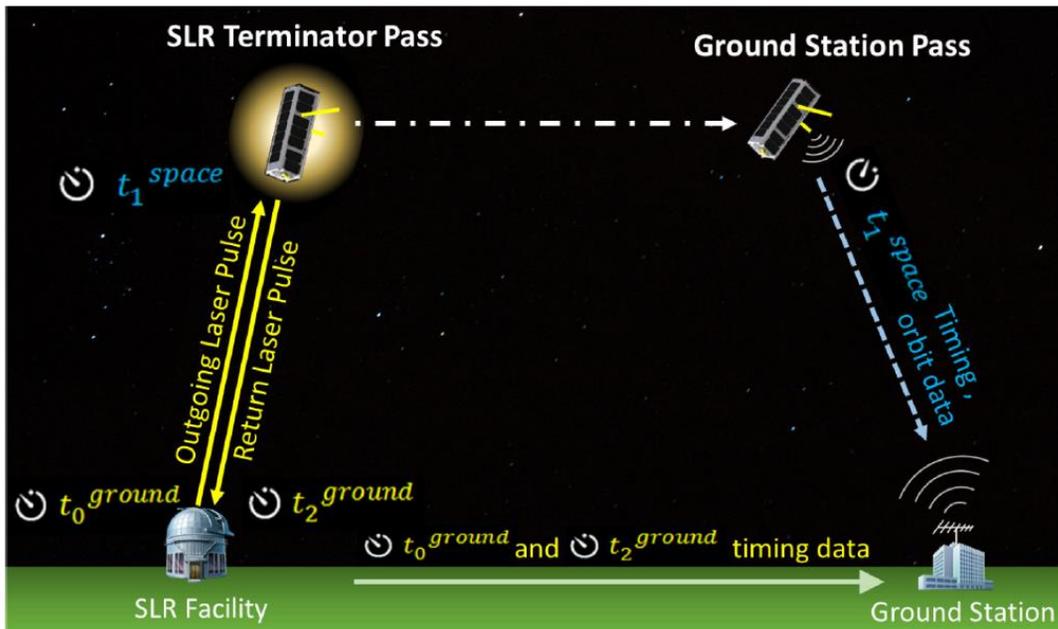
<sup>2</sup>R&M Technologies

# Objective & Motivation

- Enable NRL to participate in ILRS and other Laser Time Transfer (LTT) experiments, including NASA's CHOMPTT and ESA's ACES/ELT
  - leverage COTs equipment and technological advances from ILRS to date
- ACES/ELT
  - launch to ISS: expected (2020?)
  - ultra-stable atomic (Cs fountain + H-maser) clock ensemble<sup>1,2,3</sup>
  - microwave link<sup>10</sup> for ACES primary time transfer mode
  - 532nm laser link<sup>12</sup> for optical timing experiments
    - gated detector: laser pulse on target within 100ns
- Objectives of optical link payload:
  - evaluate limits in comparing precision ground clocks via LTT utilizing ACES timescale
  - improve atmospheric propagation models by comparing refractive index to microwave propagation delay
  - optically derived precision orbits for ISS

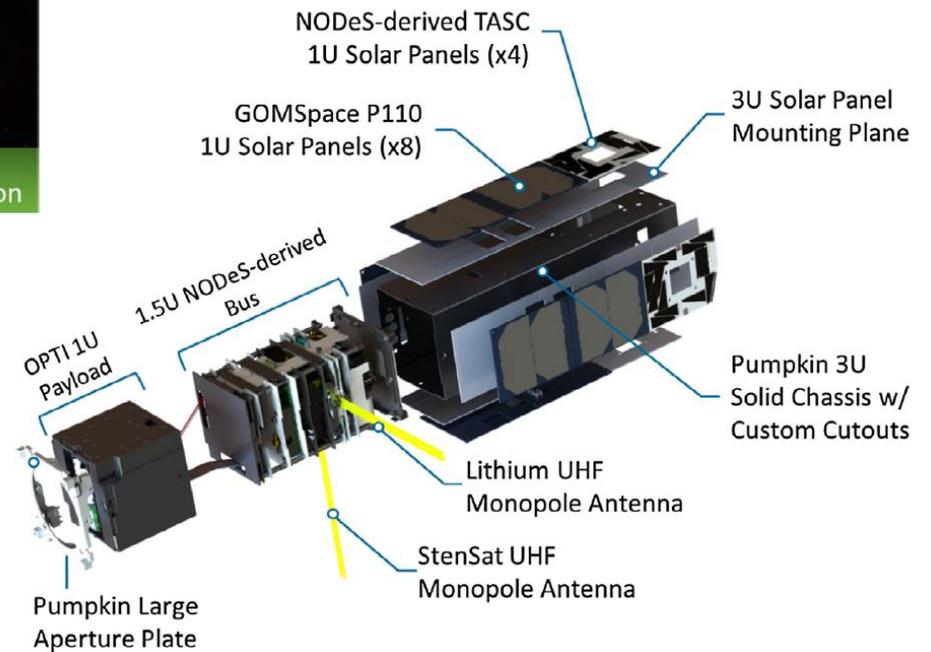


Source: <http://www.esa.int>



- NASA Ames (bus)
- Univ. of Florida (OPTI)
- NRL & Univ. of Florida ground stations
- Launch expected late 2018

1U OPTI Payload



Anderson et al., in press (Adv. Space Res.)

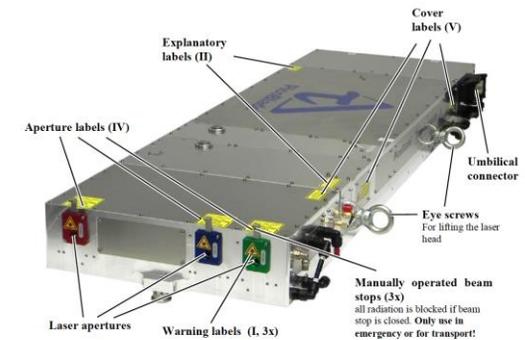
## Telescope: Brashear 1 meter telescope

- All reflective design
- F#: 89
- Focal plane: @12.640 m
- Slew rates:
  - 15 degrees/sec slew rate (elevation)
  - 25 degrees/sec slew rate (azimuth)
- Pointing accuracy: <2 arcsec RMS all sky

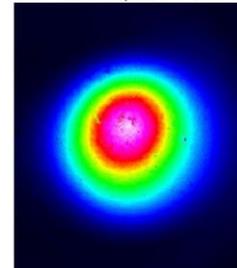


## Laser: Lumentum PicoBlade

- Ultra-short pulses, passively stabilized
  - ~28 ps (532 nm)
  - ~34 ps (1064 nm)
  - Single-shot to 20 kHz capable
  - 82 MHz oscillator (syncs to high precision external clock)



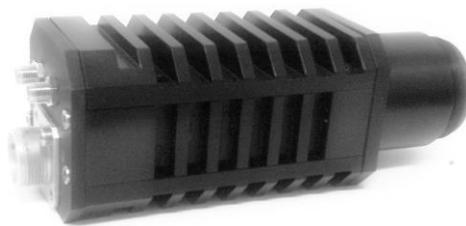
$\lambda = 532 \text{ nm}$   
PRF = 1.5 kHz, 1000 mm from exit window



Repetition rate (kHz)	Measured power (W) $\lambda = 532 \text{ nm}$	Pulse Energy ( $\mu\text{J}$ ) $\lambda = 532 \text{ nm}$
20	5.85	292.6
10	4.57	457.0
5	3.14	627.0
3	2.04	680.8
2.5	1.73	691.6
2	1.39	693.5
1.5	1.04	693.5
1	0.90	717.8

## Rx detector: Compensated Single Photon Avalanche Detector (C-SPAD)

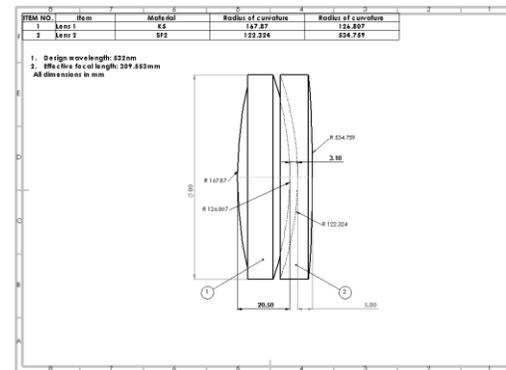
- Si APD
- 200  $\mu\text{m}$  active area
- Quantum Efficiency: 40%
- AR coated for 532
- Accepts 12 mm diam beam
- FOV: 1 degree
- Active quenching circuit
  - Time walk compensation  $< \pm 10$  ps



PESO Consultant Ltd.

## Optical train:

- Custom optical elements were designed at NRL for better and efficient coupling of the laser system into the telescope
- High quality optics (mirror, polarizers, lenses) were acquired for system efficiency



## Detector gating: Graz Range Gate Generator

- medium Resolution Event Timer and range gate generator
- 5ns resolution in time stamping
- 500ps resolution programable range gate generator
- accurate enough for generating range gates, range residuals, and real time plots for displays



## Event timing: New Picosecond Event Timer (NPET)

- supports 2kHz epoch timestamping
- <math><0.9\text{ps}</math> timing jitter per channel
- <math><0.5\text{ps}</math> timing drift per Kelvin
- <math><0.1\text{ps/hour}</math> timing stability
- requires spectrally clean clock signal



**PESO Consultant Ltd.**

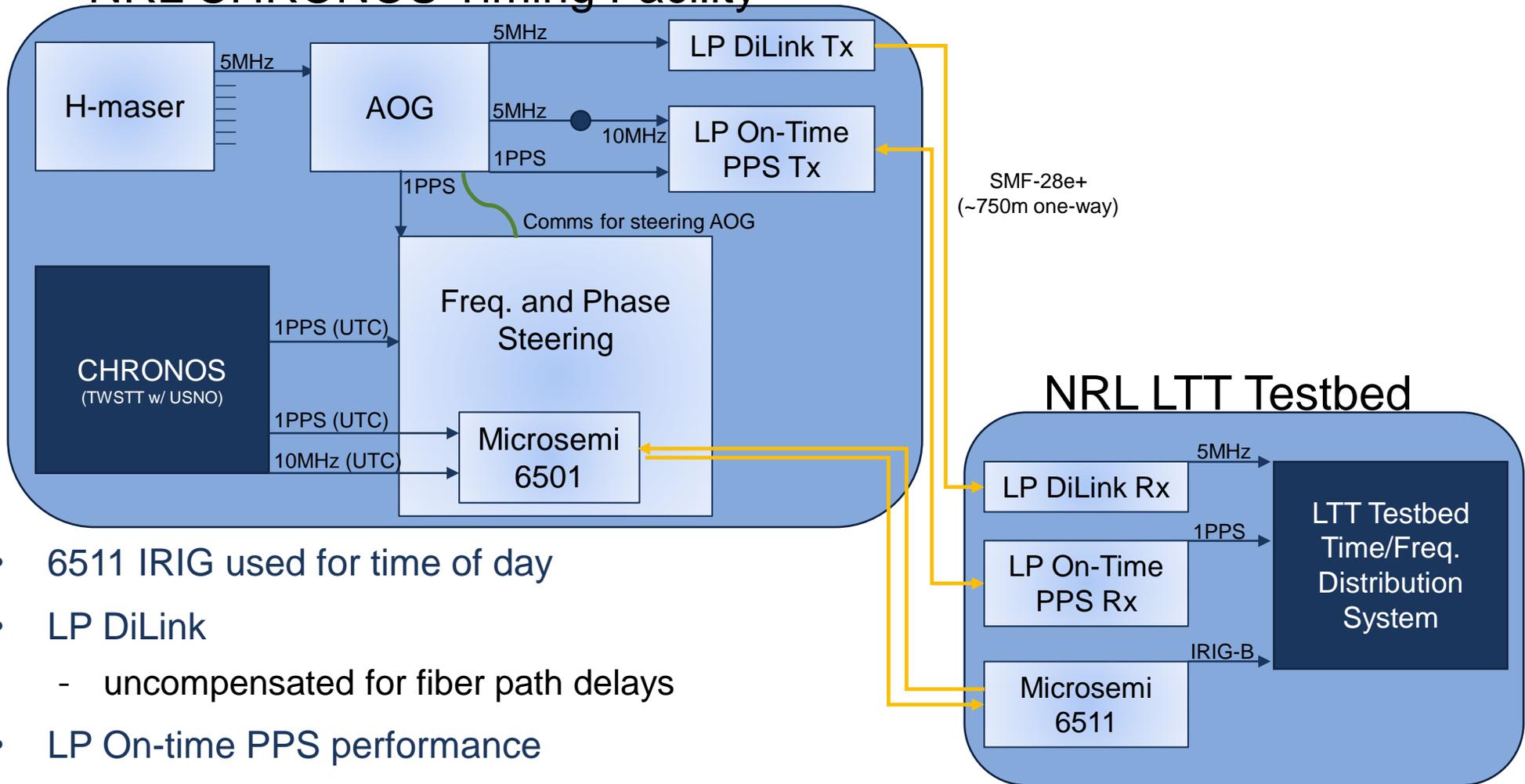
# New Reference Timing Signals

- High precision atomic frequency reference
  - Microsemi H-maser w/ LPN option
  - AOG used for steering to UTC
  - Dedicated SMF-28e+ host <-> LTT testbed
    - ~750m one-way
  - Microsemi 6511 (coarse time-of-day)
    - TWTFT over fiber with CHRONOS 6501
    - <20ns performance
  - Linear Photonics On-time PPS to LTT testbed
    - PPS time marker aligned to UTC(USNO)
    - 276ps ( $\pm <1$ ps) static offset Tx/Rx
  - Linear Photonics DiLink to LTT testbed
    - delivers H-maser 5MHz frequency reference
    - Uncompensated for fiber delay variations
- Initial integration complete August, 2018
- Ongoing monitoring of signals at H-maser and LTT testbed



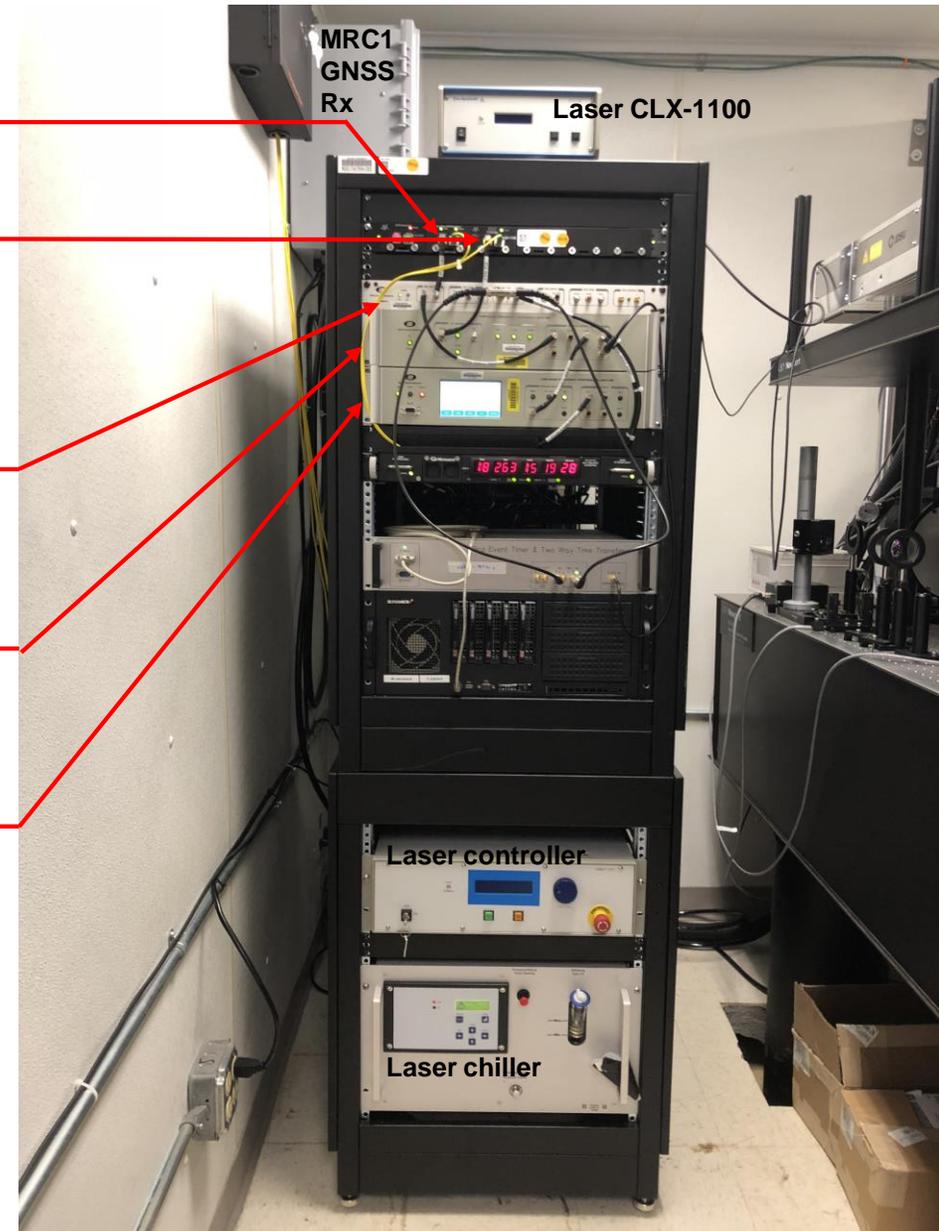
AOG-110 Auxiliary Output Generator

## NRL CHRONOS Timing Facility

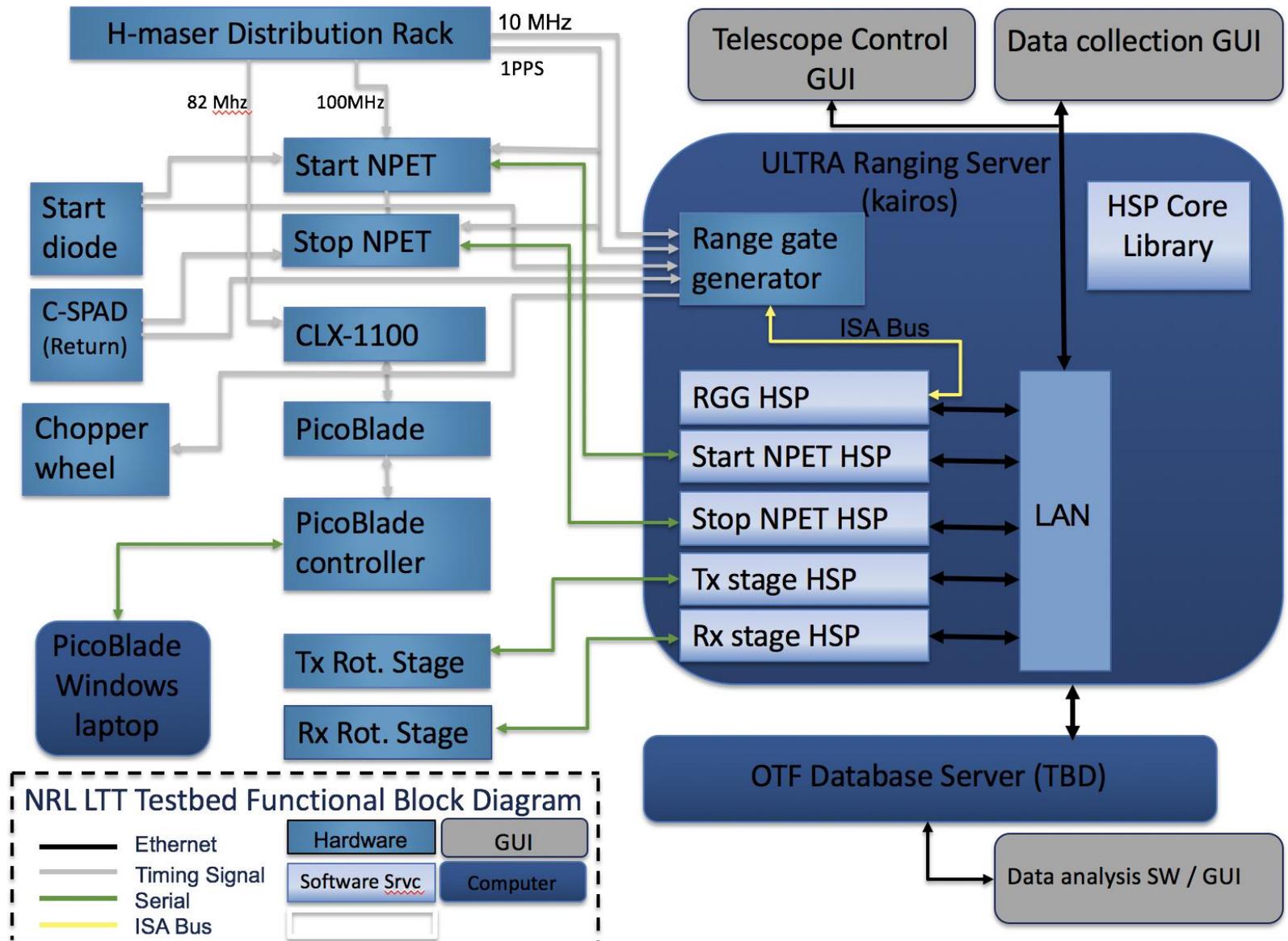


- 6511 IRIG used for time of day
- LP DiLink
  - uncompensated for fiber path delays
- LP On-time PPS performance
  - Tx – Rx offset: 276ps
  - over ~1500m roundtrip

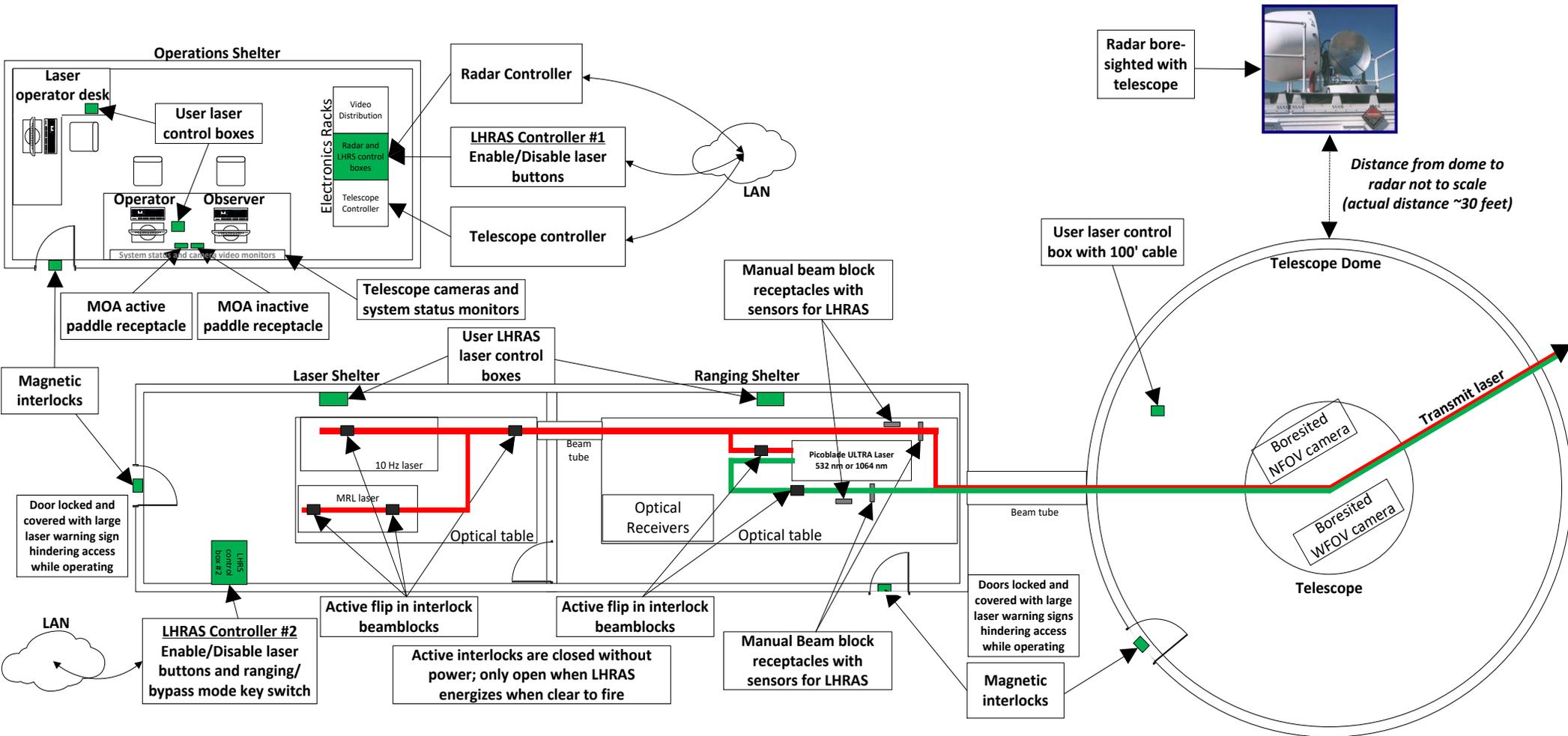
- Linear Photonics Timelink
  - On-time PPS Rx
    - 276ps ( $\pm < 1$ ps) static offset Tx/Rx
  - DiLink Rx
  
- Spectra Dynamics, Inc.
  - PPS generator
    - timing reference for event timers
    - aligned to UTC(USNO) within  $\sim 250$ ps
  - low-noise frequency cleanup osc.
    - 100MHz output for event timers
    - 10MHz output for synthesizer
  - low-noise frequency synthesizer
    - 82MHz for laser CLX-1100
    - CLX-1100 measures laser oscillator jitter wrt input 82 MHz =  $\sim 0.10$ ps
  
- Microsemi PPS and RF Amps
  - timing source for legacy SLR systems



# System Level Functional Diagram

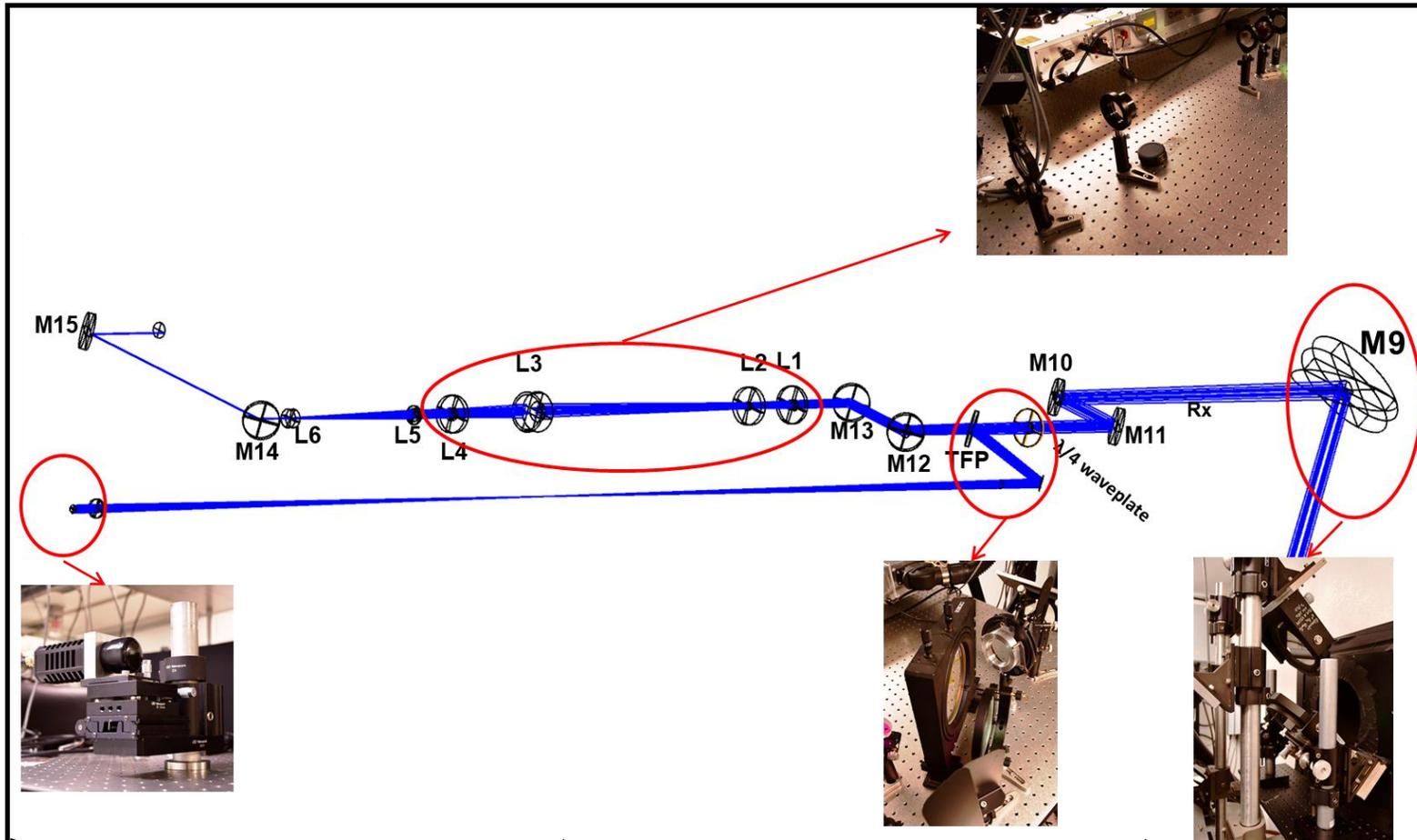


## New interlock (complete) and radar (late 2018)

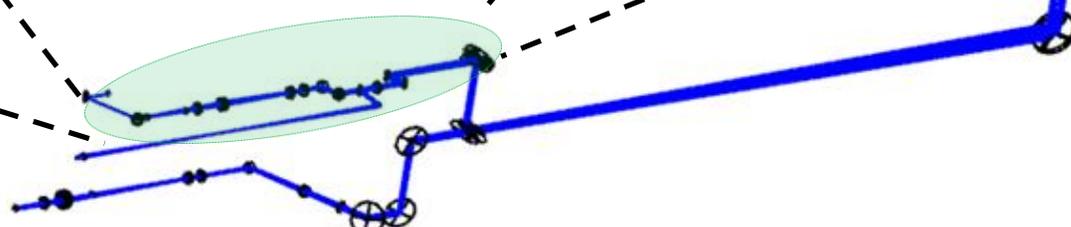
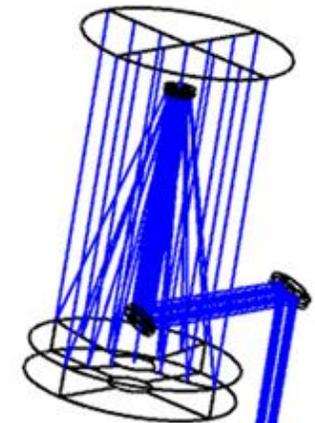


**Callout boxes indicate all components in the LHRAS**

# NRL LTT Testbed Optical Layout



- Class 1M
- $50 \mu\text{rad } 1/e^2$  HW divergence

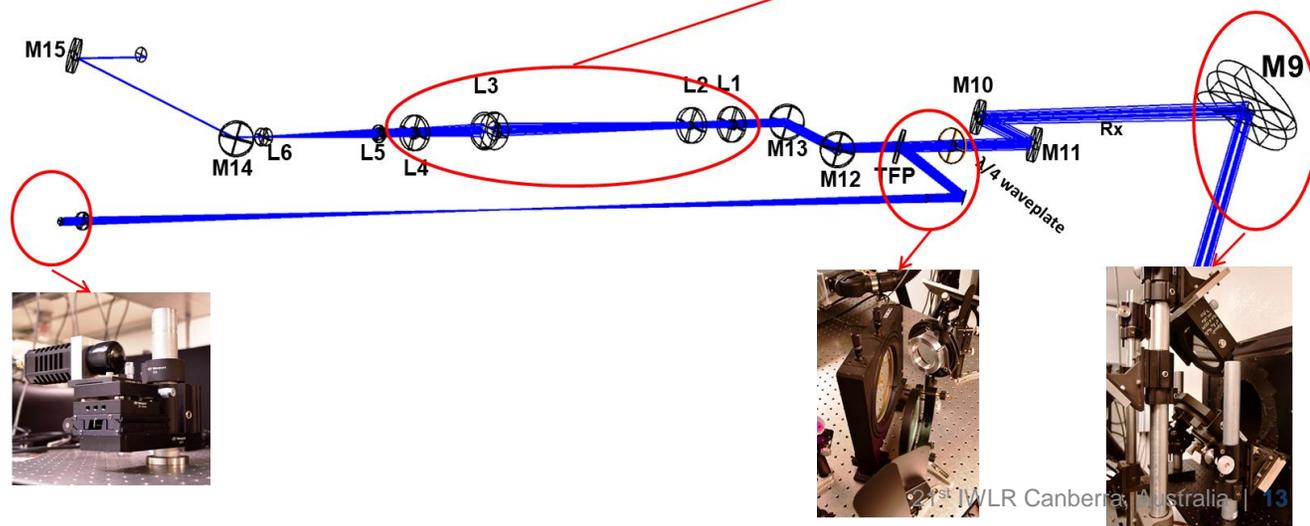
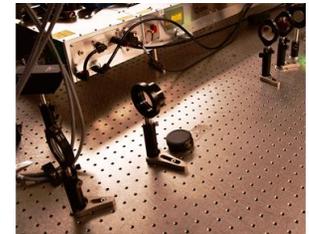


- **Transmitted beam**

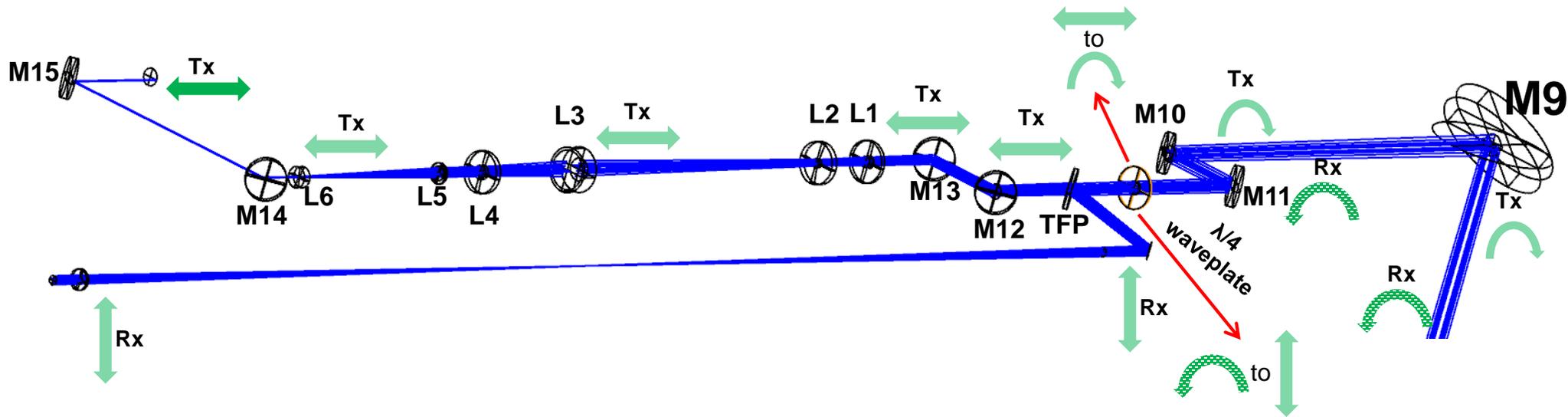
- Collimated from 1 mm to 12 mm diameter
- Optics for matching telescope F# w/ 100  $\mu$ rad FW divergence
- Folding mirrors for fine alignment with telescope
- Polarizing optics to separate Tx signal from Rx signal
- Periscope insertion into telescope optics

- **Received beam**

- Thin film polarizer splits return light into Rx arm path
- Collimated to 12 mm diameter
- Return photons directed and aligned onto Rx detector (C-SPAD)



Initial design for separating Tx and Rx optical beam by polarization components



Polarization components measured at detector position:

- 85% vertical
- 7% Horizontal

- Test Performed to the system:

- Initial collimation: **GOOD**
- Optical elements where tested at NRL to characterize their optical performance:  
**Current loss: ~30% (will be improved by fixing collimation size)**

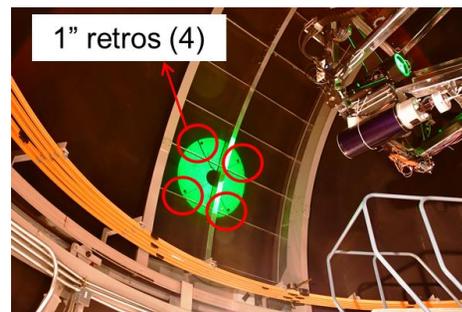
**Total expected loss: ~ 15 to 20% (transmitting)**

- System alignment with telescope: **GOOD**
- Polarization states maintain through the system: **GOOD**  
**92% linear polarization at the detector**
- Receiving arm focusing efficiency: **GOOD**
- System backreflections: **OK**, except >1 nJ from telescope covers (sun avoidance)

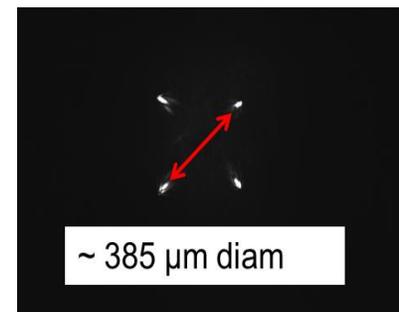
- Tests to be performed:

- Rx effective FOV
- Collimation out of the telescope

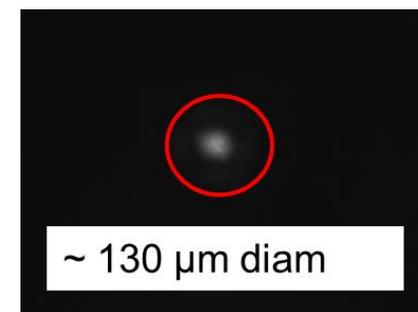
Receiving arm test, using a 75 mm fl lens



At the telescope



5.3 mm from  
system focus



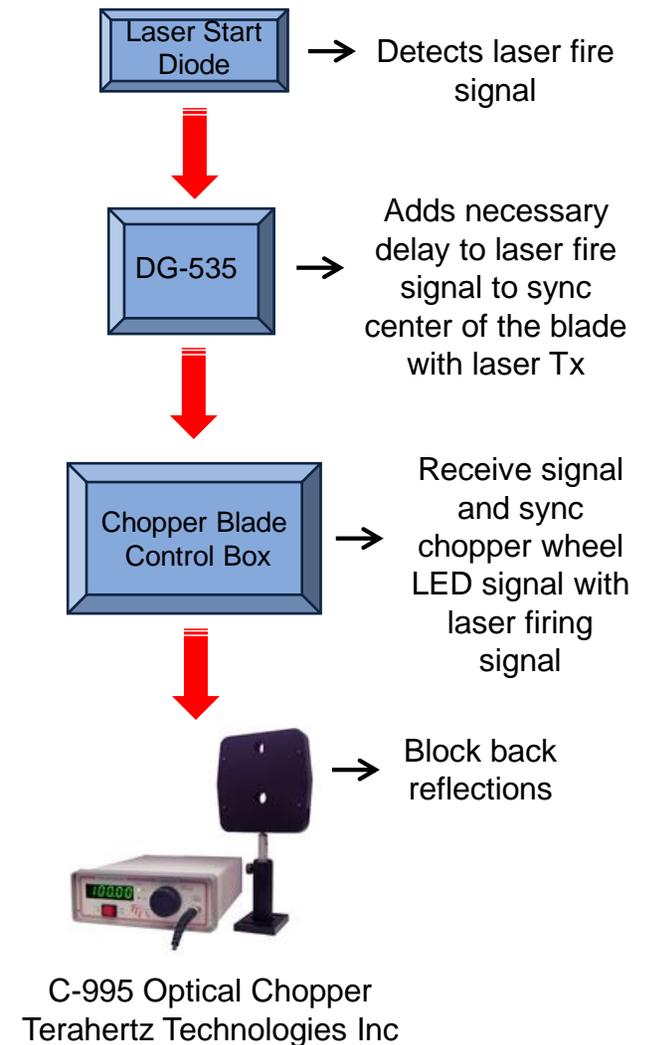
At system focus

# Backscatter Suppression

- Optical Chopper Blade implemented to suppress on-axis backscatter
- Custom blade designed to maximize opening for Rx signal, and protecting detector while Tx w/ gate open
- Located at focal plane on Rx leg
- Custom design:
  - sync with laser fire while modifying blocking duty cycle
  - 16% duty cycle, blocks 16  $\mu$ secs while sync @ 1kHz,
  - blades sized to block area of detector while adding enough buffer to keep protecting in case of signal jitter
  - tested to sync up to 1.5 kHz
  - Inner and outer ring made to maximize opening duty cycle

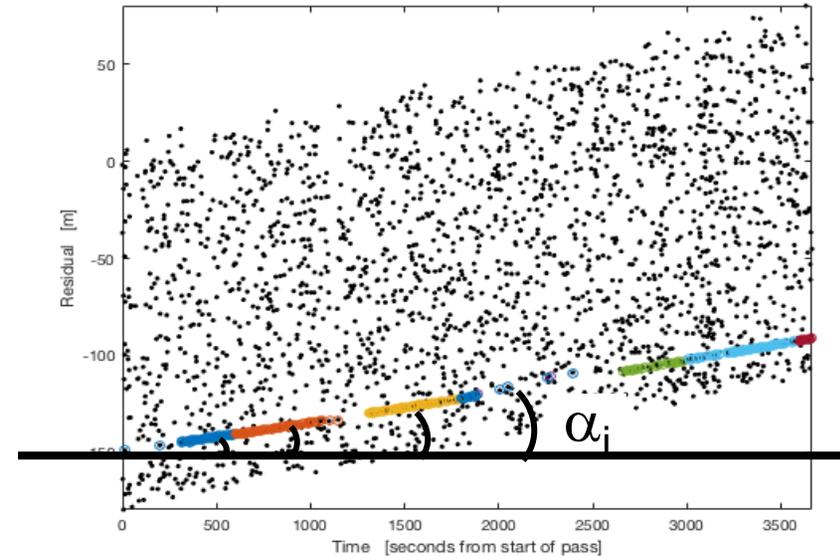


Custom optical chopper blade, designed and built at NRL



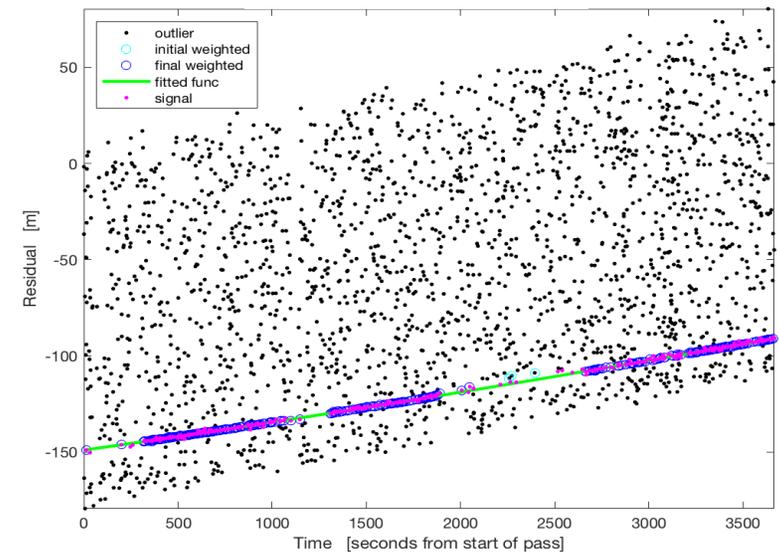
## Step #1: Initial search

- Two-modes: ground calibration and satellite
- For each 10-min interval,  $i$ ,
  - determine direction,  $\alpha_i$ , that minimizes width of a high-resolution histogram and contains bin with maximal number of data points
  - select SLR measurements for residuals that fall within narrow band along direction  $\alpha_i$ 
    - thickness of the band is a function of system jitter and target signature

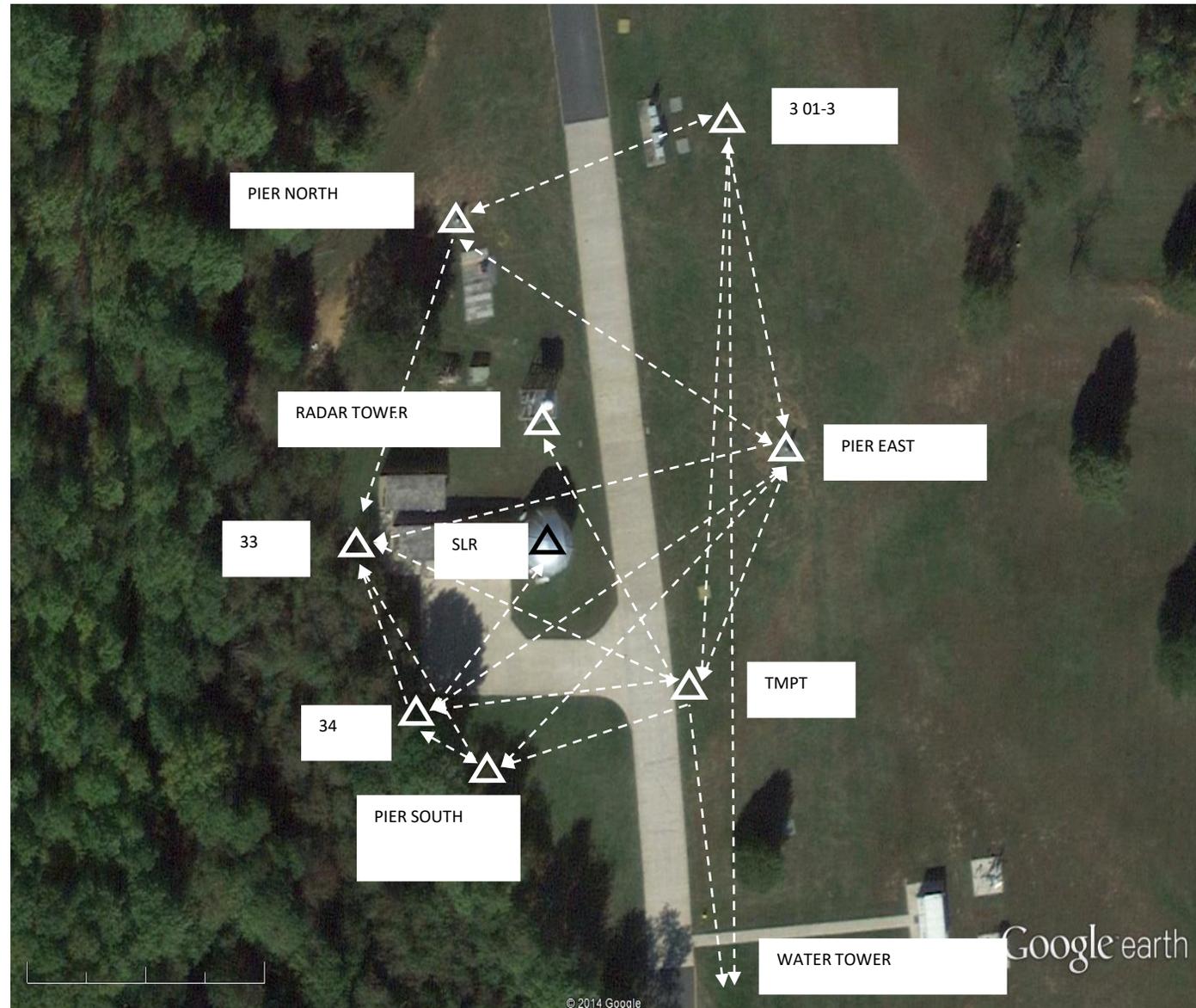


## Step #2: Outlier rejection

- Iterative weighted least-squares of regression function to find signal photons
  - all data points are included
  - data found in Step 1 (cyan) used for initial weighting
  - subset of initial weighted data points remain (blue) after iterative fitting and outlier rejection
  - solution converges when no outliers remain
  - full-rate signal photons (magenta) are all remaining data points

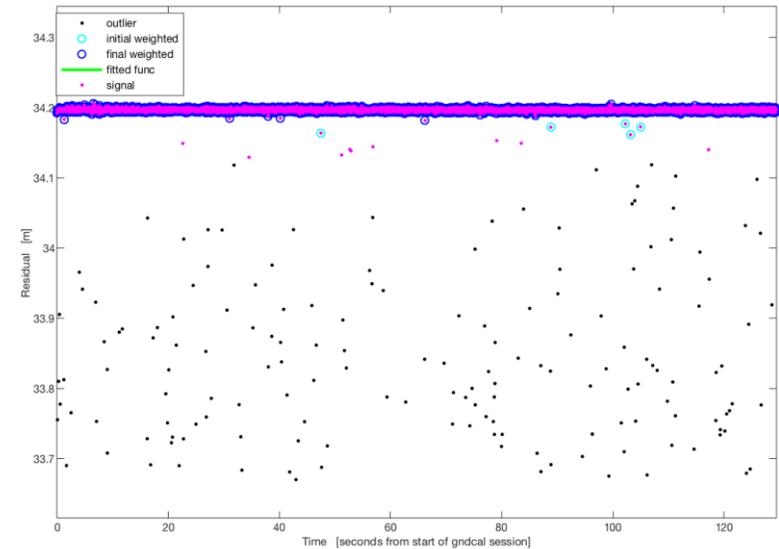


- Completed in 2016 by NOAA NGS
- Tie between ground ranging targets and NRL telescope realized via AXIS software (Geoscience Australia)





peas (16 Oct 2018)



delay [ps]: 114068.0

drift [ps/s]: -0.008

Compares to calculated nominal delay based on

- optical path length (zemax)
- measured cable delays
- electronic signal rise times

Validate stability (in prog)

# Summary & Next Steps

- Aim to enable NRL participation in ILRS and other LTT experiments
  - initial effort engineering to design 532nm system, requirements driven by ACES/ELT
  - study potential new methods to characterize, monitor, compensate for system delays
  - evaluate limits of LTT technique
- Initial 532nm optical layout designed and integrated
  - will test polarization-based attenuation on Tx and Rx legs for controlling flux on C-SPAD
  - need to verify 100  $\mu$ rad divergence
  - need to improve backscatter suppression for >1.5 kHz rates
- Initial integration of electronics and timing systems complete
  - finish testing s/w interfaces with laser, NPETs, and timing systems
  - develop technique for controlling laser fire time
- Developed a new tool for extracting full-rate signal photons
  - add simplified user interface for low-latency post-processing (and reanalysis)
  - verify accuracy of data products generated using the tools
- Calibration and validation
  - classical methods used for ongoing system checkout and characterization
  - interested to find/explore alternate methods

# Questions?