

Evaluation of PPN parameter Gamma as a test of General Relativity using SLR data

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$$\Delta \ddot{\vec{r}} = \frac{GM_E}{c^2 r^3} \left\{ [2(\beta + \gamma) \frac{GM_E}{r} - \gamma \dot{\vec{r}} \cdot \dot{\vec{r}}] \vec{r} + 2(1 + \gamma) (\dot{\vec{r}} \cdot \dot{\vec{r}}) \dot{\vec{r}} \right\} +$$

$$(1 + \gamma) \frac{GM_E}{c^2 r^3} \left[\frac{3}{r^2} (\dot{\vec{r}} \times \dot{\vec{r}}) (\dot{\vec{r}} \cdot \vec{J}) + (\dot{\vec{r}} \times \vec{J}) \right] +$$

$$\left\{ (1 + 2\gamma) \left[\ddot{\vec{R}} \times \left(\frac{-GM_S \ddot{\vec{R}}}{c^2 R^3} \right) \right] \times \dot{\vec{r}} \right\},$$

c = speed of light,

β, γ = PPN parameters equal to 1 in General Relativity,

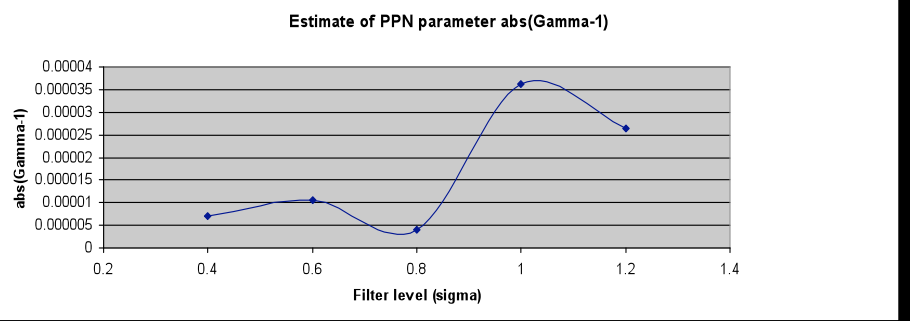
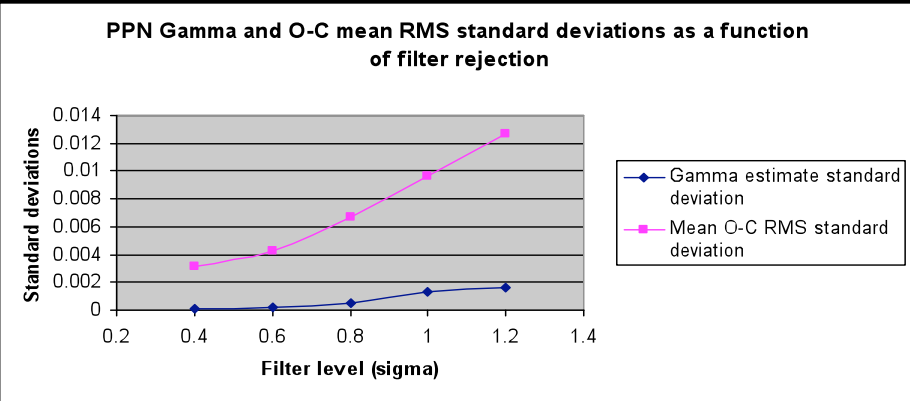
$\dot{\vec{r}}$ is the position of the satellite with respect to the Earth,

$\ddot{\vec{R}}$ is the position of the Earth with respect to the Sun,

\vec{J} is the Earth's angular momentum per unit mass
 ($|\vec{J}| \cong 9.8 \times 10^8 \text{m}^2/\text{s}$), and

GM_E and GM_S are the gravitational coefficients of the Earth and Sun, respectively.

- Gamma estimated as a solve-for parameter in least-squares sense
- RMS of O-C used as criteria for filter to reduce effect of mismodelling
- SLR strength is 2-way range, and GR radial acceleration is mostly radial
- Results preliminary, longer time span, combination of LAGEOS/LAGEOS II and evaluation of different gravity models to follow, effects of modelling improvement should be apparent in errors



Filter (σ)	PPN $ \bar{\gamma} - 1 $	σ	Mean O-C RMS (m)	σ (m)
0.4	6.977×10^{-6}	0.000078014	0.005403	0.003136
0.6	1.0649×10^{-5}	0.000183666	0.007469	0.004251
0.8	3.947×10^{-6}	0.000525783	0.011888	0.006701
1.0	3.268×10^{-5}	0.001324442	0.017238	0.009609
1.2	2.6428×10^{-5}	0.001647755	0.023074	0.012702